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Calculation of opacities and emisivities for carbon plasmas under NLTE and LTE conditions

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Abstract. We calculate different optical properties for carbon plasma in a wide range of temperatures and densities by using ATOM3R-OP code which has been recently developed. We have calculated average ionizations, level populations, opacities and emissivities and we focus our study on the identification with our code of Coronal Equilibrium, Non Local Thermodynamic Equilibrium and Local Thermodynamic Equilibrium regions for this kind of plasma. Moreover, we analyse the differences in the optical properties when they are calculated in Non Local Thermodynamic Equilibrium and Local Thermodynamic Equilibrium.

1. INTRODUCTION

ATOM3R-OP is a code recently developed to calculate optical properties for a wide range of temperatures and electron number densities. In this code we have implemented the rate equations, the Saha equation (LTE) and Coronal Equilibrium model (CE), and in a previous work [1] it has been shown that the ionic populations obtained by solving rate equations are in good agreement with those prodicted by LTE and CE at high and low electron density respectively.

In this work we make an study of optically thin hot carbon plasma for a wide range of temperatures and densities. First of all, we focus our study on the identification of the CE, NLTE and LTE regions for this kind of plasma by comparison of the average ionization and level populations. Secondly, we calculate the opacities and emissivities and we show the differences observed when the ionic populations are calculated using LTE, CE and rate equations. The atomic data used in this work are obtained from the FAC code [2] which provides relativistic energies, bound-bound oscillator strengths and bound-free cross sections.

2. DESCRIPTION OF THE CALCULATIONS

2.1 ATOM3R-OP CODE

ATOM3R-OP is a code to calculate optical properties for a wide range of temperatures and densities. This code obtains ionic populations by solving the steady state rate equations for optically thin and thick plasmas. For optically thick plasmas we consider homogeneus plasmas in planar, cylindrical and spherical geometries and the escape factor formalism is used. The opacities and emissivities are obtained

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using Voigt profile for all spectral lines and assuming complete redistribution of the photons. The atomic data necessary to calculate the quantities mentioned above are obtained from different atomic models, i.e. selfconsistent or analytical models and from atomic databases. A large number of ionic configurations can be considered using sparse matrix techniques to store the matrix elements and iterative methods to solve the rate equations.

Table 1. Atomic Data.

	Ionization Potential (eV)	Levels Number	B-B Transitions Number	B-F Transitions Number
CI	11.2603	1285	34870	277
CII	24.3833	104	744	2523
CIII	47.8879	486	7720	4923
CIV	64.4939	224	2565	3566
CV	392.0872	449	6806	10087
CVI	489.9933	225	2621	169

2.2 CARBON ATOMIC DATA

The atomic data used in this work are obtained from FAC code which provides relativistic energies, bound-bound oscillator strengths and bound-free cross sections with a convenient specification of mixing schemes and it also includes configuration mixing. Some details of the atomic data are shown in the table 1. Finally, Krammer's formula is used for the free-free cross section and Voigt profile is used for all the spectral lines.

3. RESULTS AND CONCLUSIONS

We have determinated the average ionization of optically thin carbon plasmas in a range of temperatures and electron number densities given by $(10-200 \text{ eV}, 10^{14}-10^{24} \text{ cm}^{-3})$. In Figure 1 we show a map which allows to identify the density and temperature regions where our model predicts CE, LTE or NLTE. To establish these regions we have proceeded as it follows. If the average ionization calculated from CE/LTE

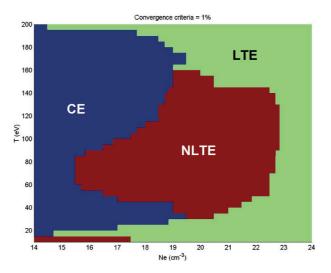


Figure 1. Map of CE, LTE and NLTE regimes determined from our level populations module.

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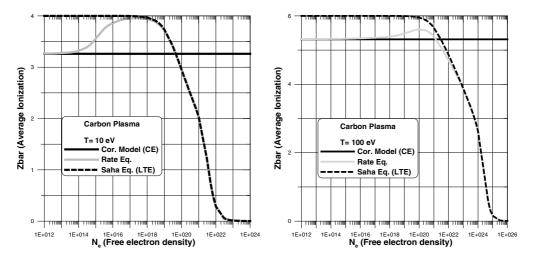


Figure 2. Average ionization vs free electron density calculated from steady state rate equation, Saha equations and Coronal Model Equations.

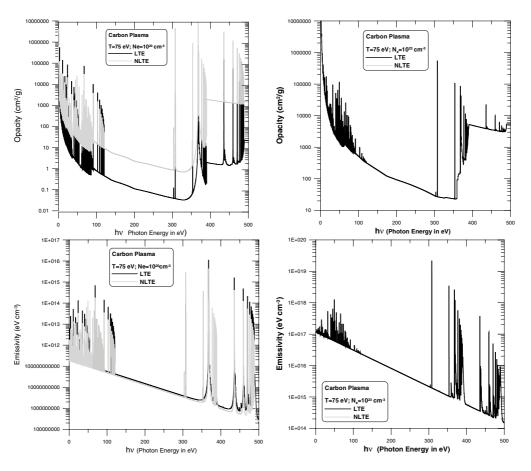


Figure 3. Opacities and emissivities for carbon plasma calculated in LTE and NLTE.

is greater than 1% respect to that one obtained from rate equations then we consider NLTE regime. Else, we assume that CE or LTE has been achieved. Detailed information about this map is shown in Figure 2.

In Figure 3 we show the multifrequency opacities and emissivities associated to two significant points (temperature and density) of the map given by Figure 1, one at NLTE (75 eV and $10^{20} \, \text{cm}^{-3}$) and other at LTE (75 eV and $10^{23} \, \text{cm}^{-3}$).

Acknowledgments

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